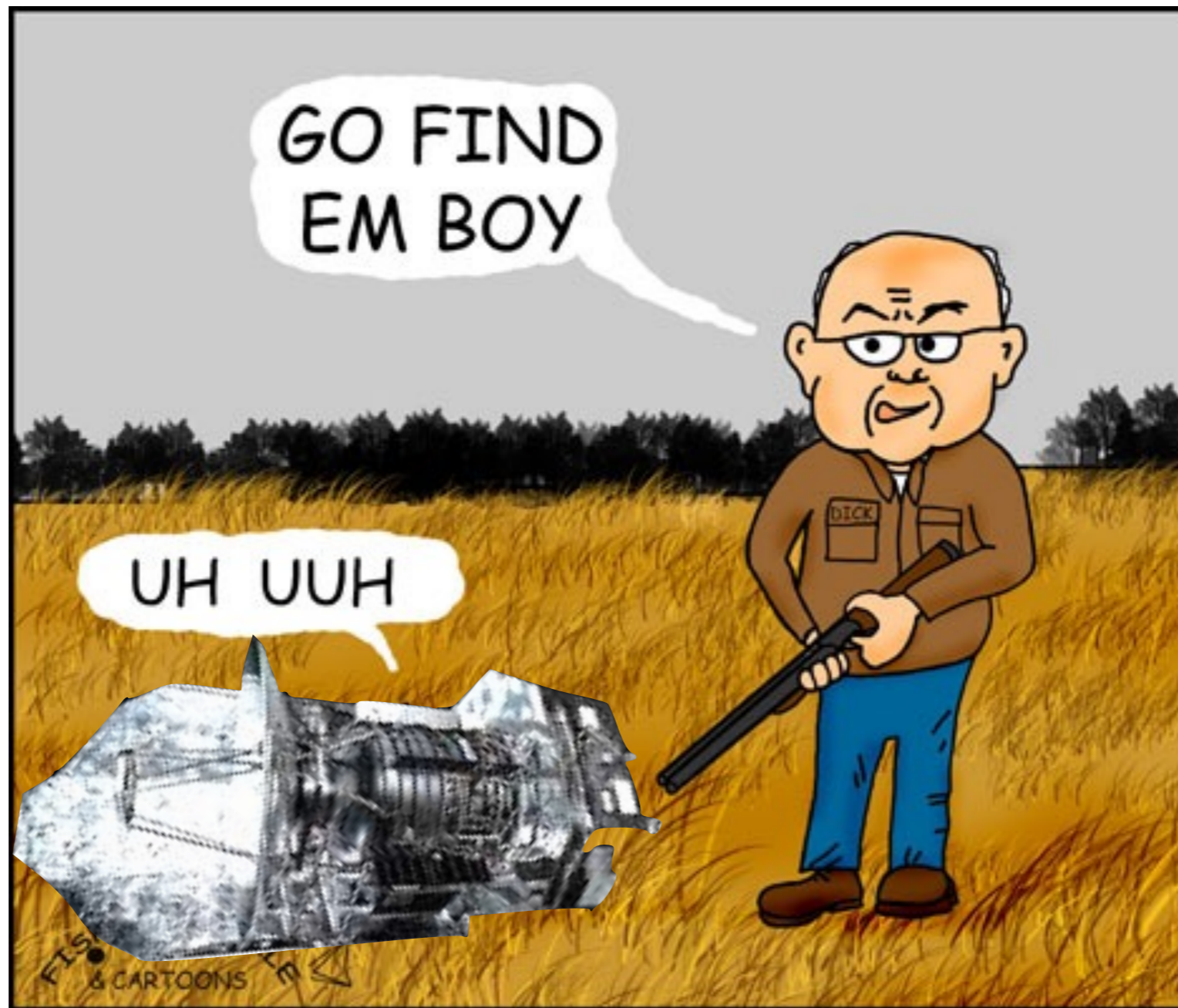
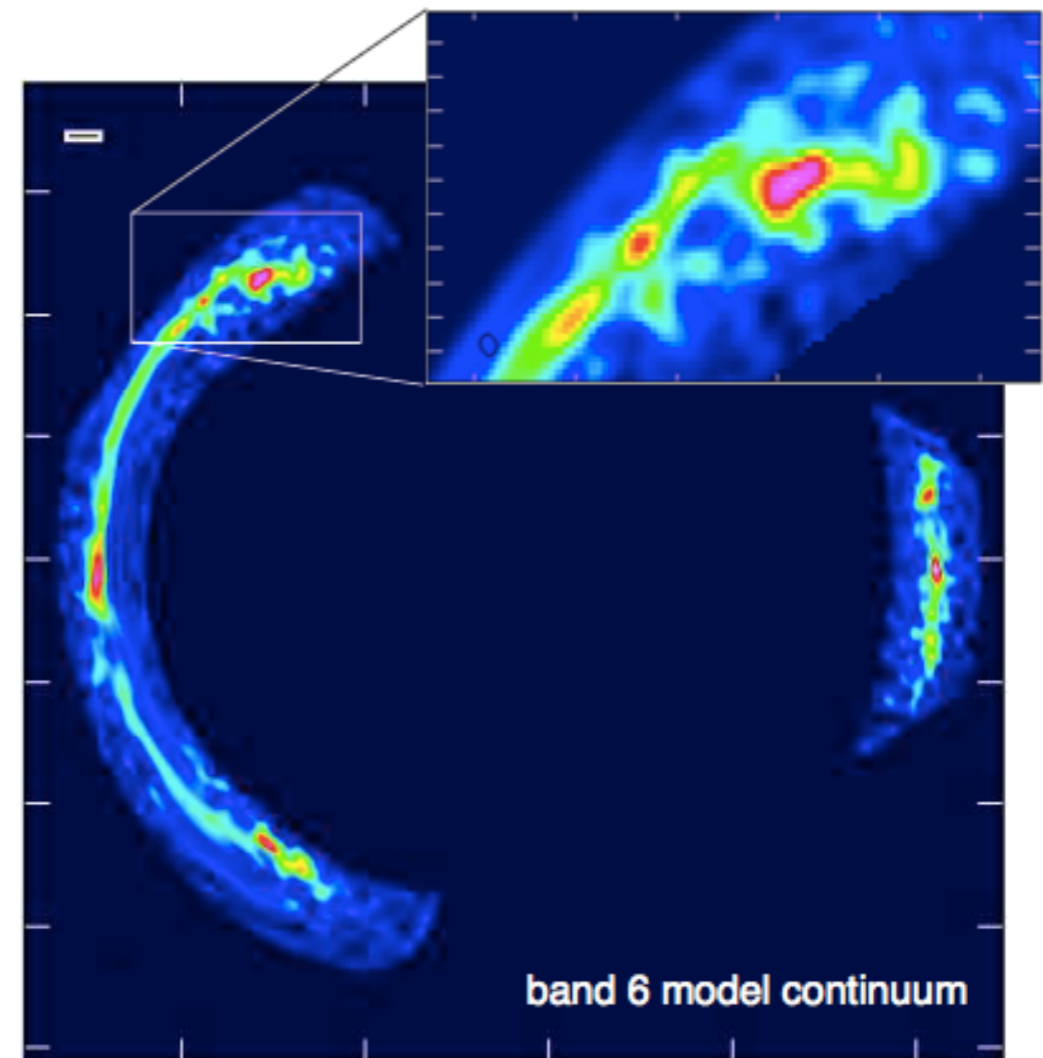


Lens hunting with *Herschel*

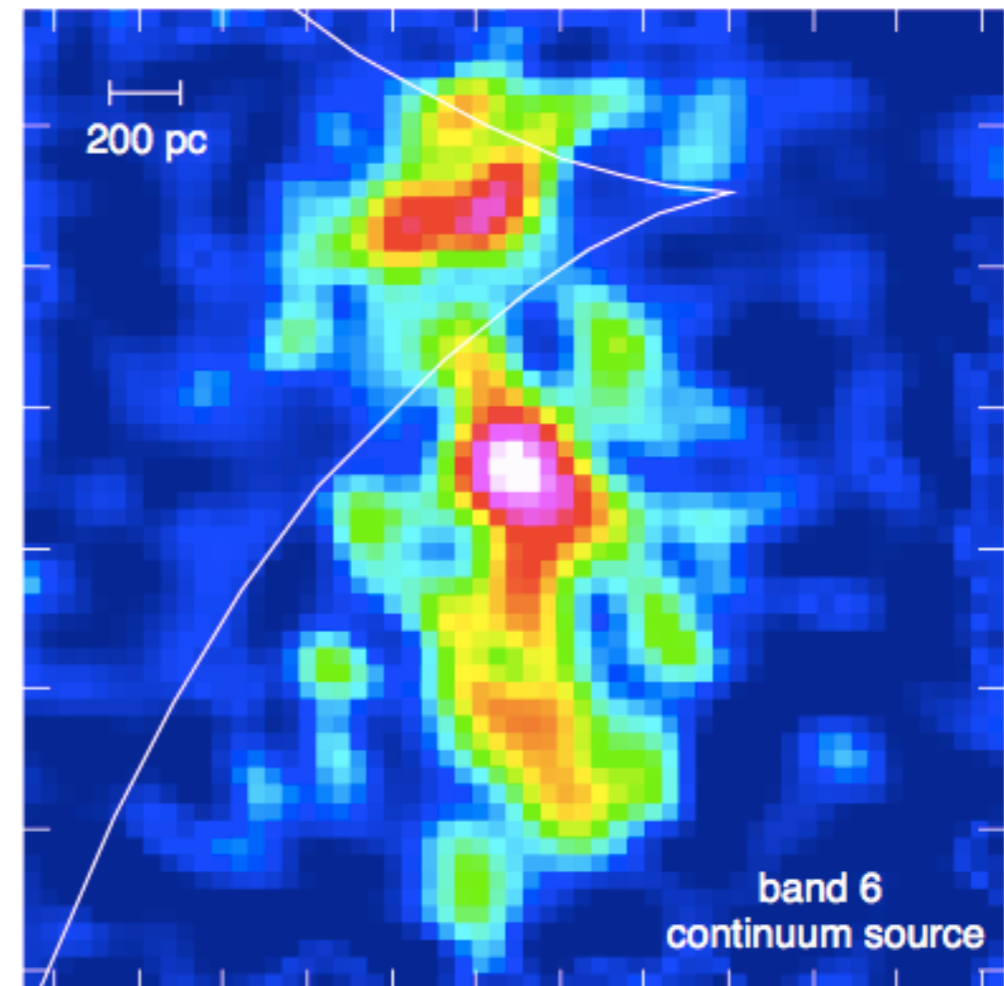
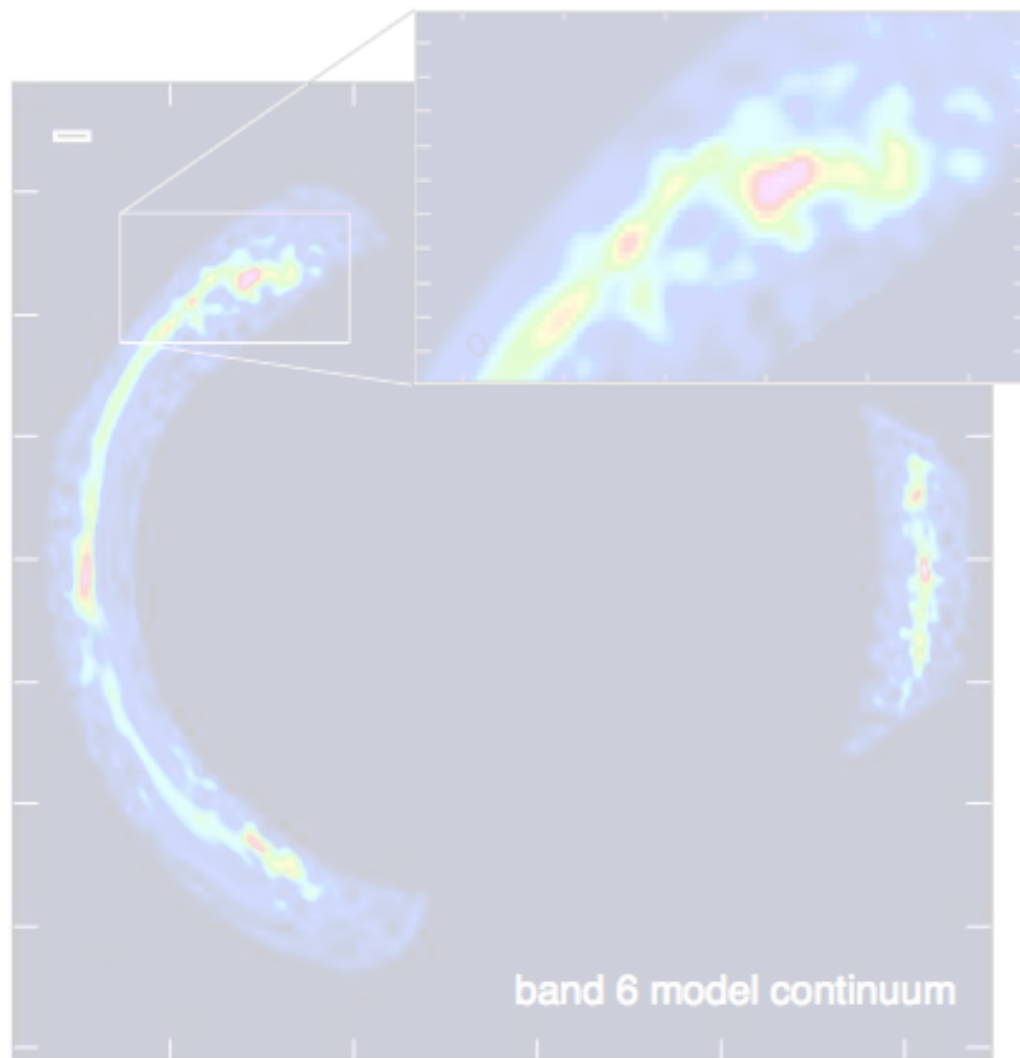
Tom Bakx



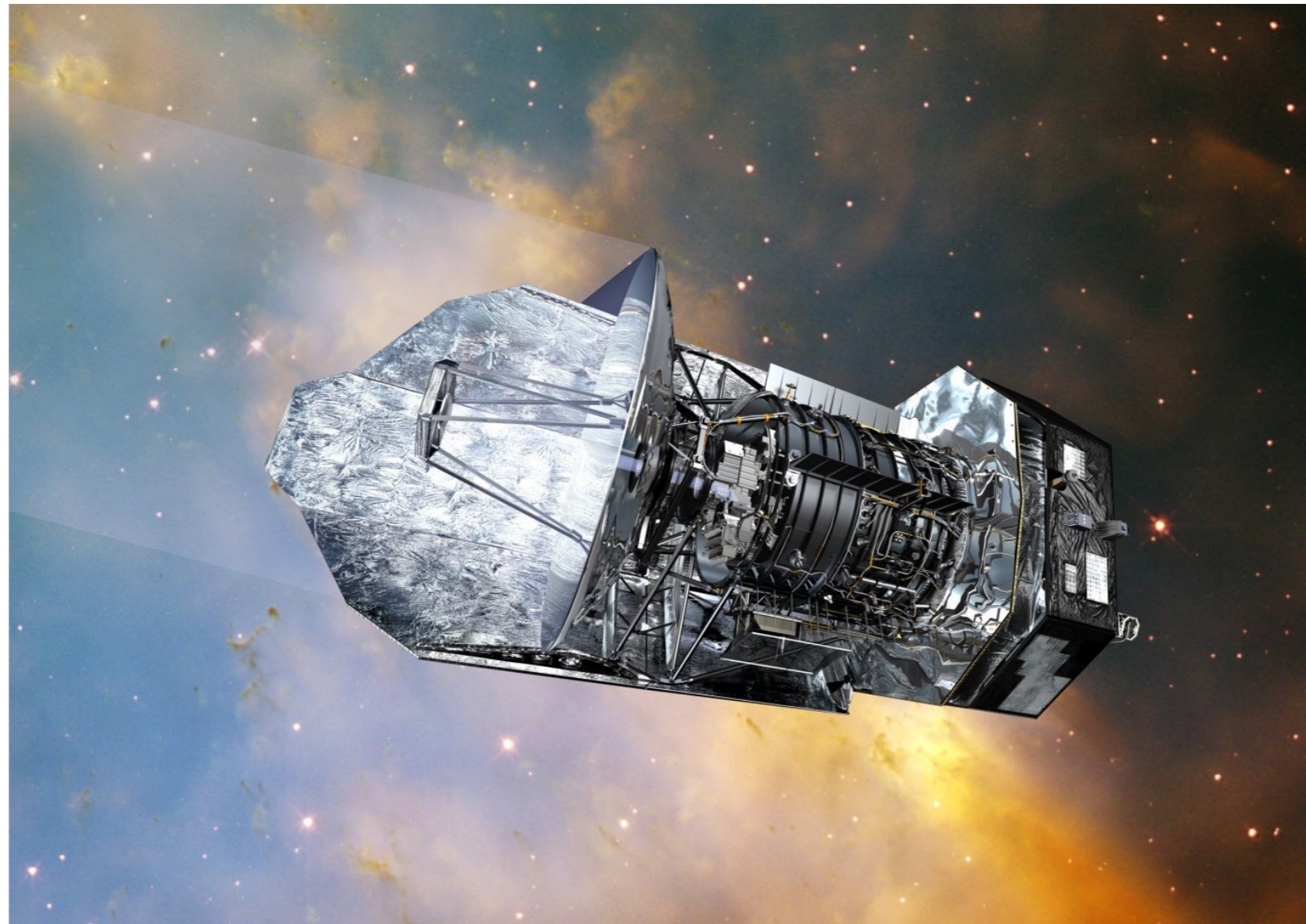
Gravitational lenses provide otherwise unreachable levels of detail



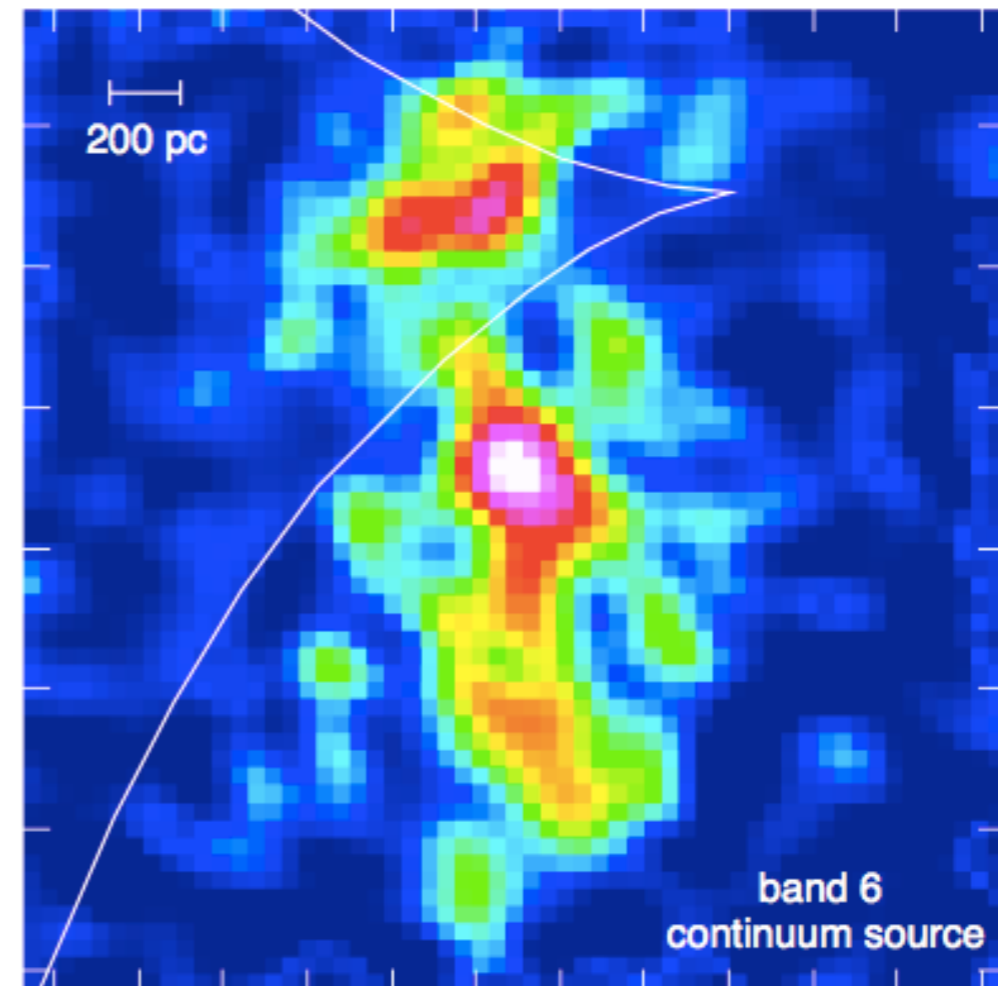
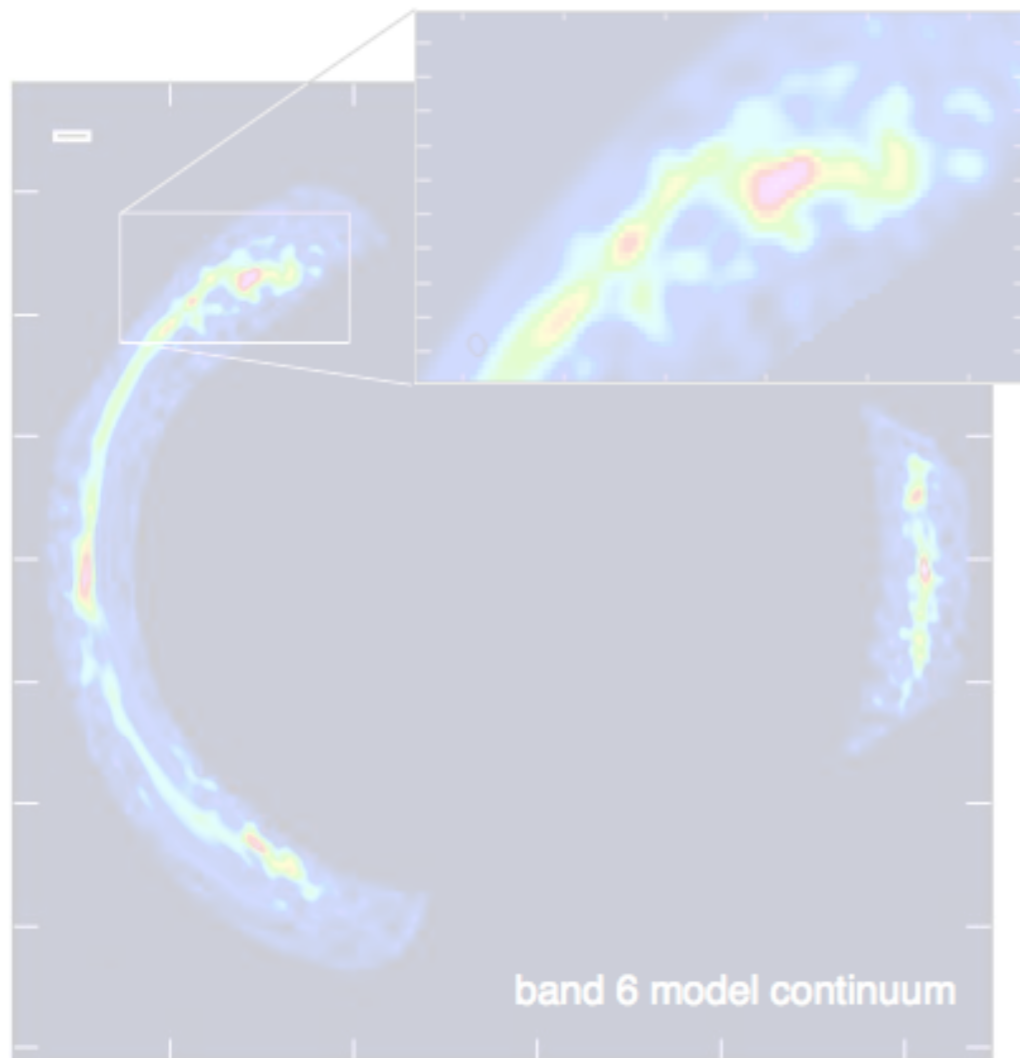
Gravitational lenses provide otherwise unreachable levels of detail



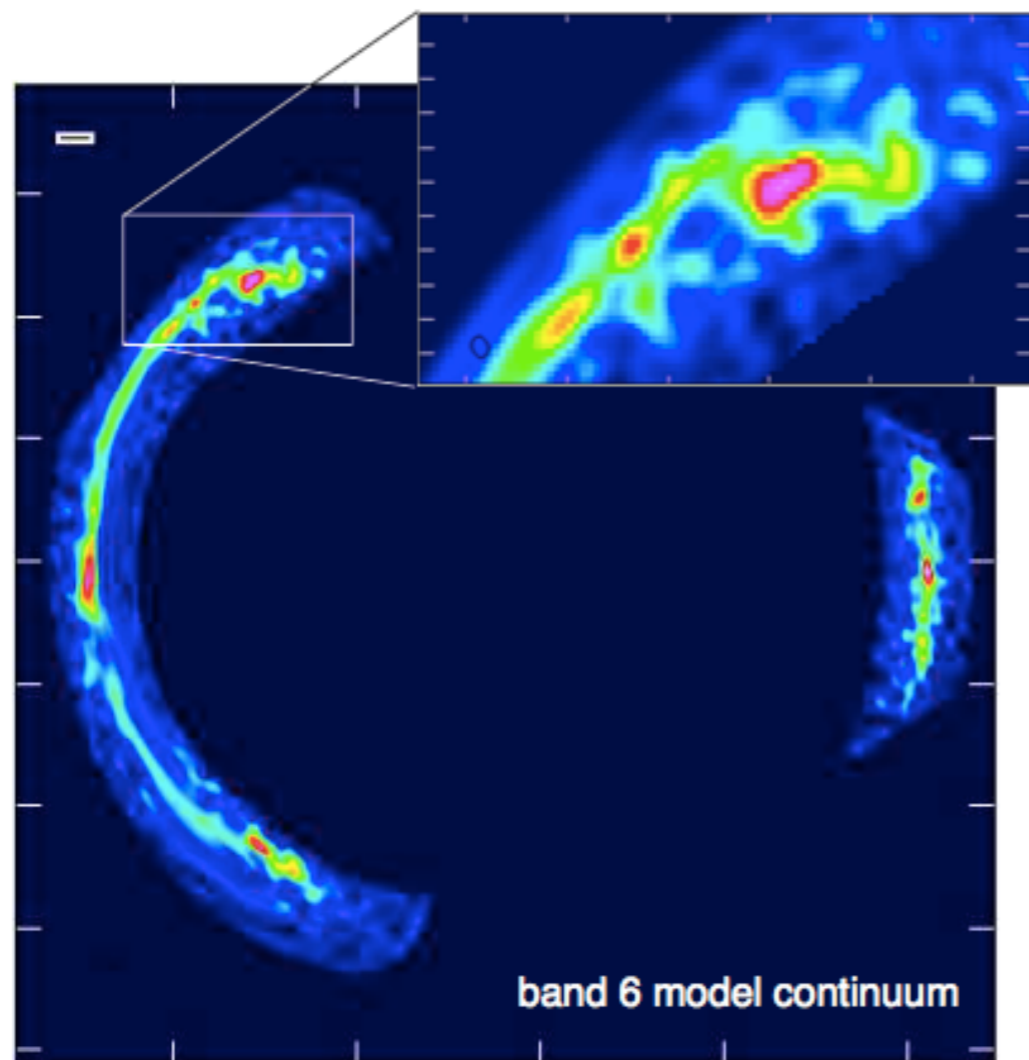
The *H-ATLAS* survey is used to find large samples of lenses



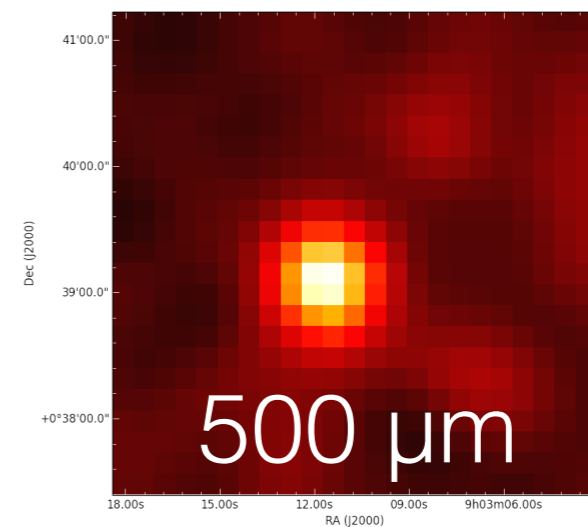
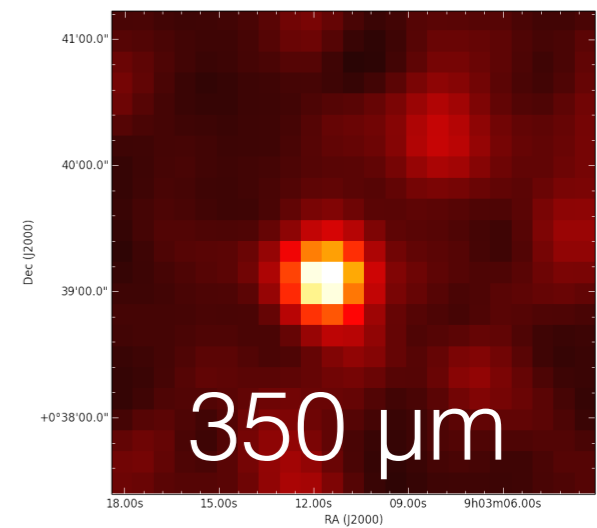
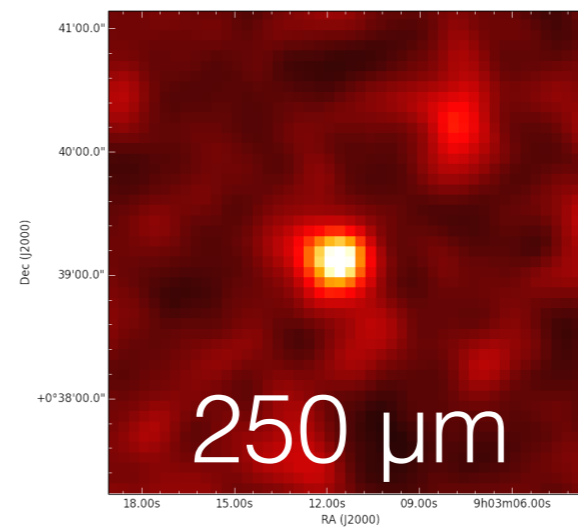
Gravitational lenses provide otherwise unreachable levels of detail



Large surveys don't have the detail of ALMA observations


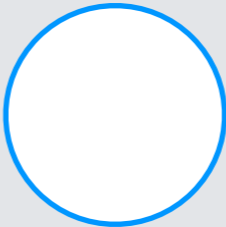
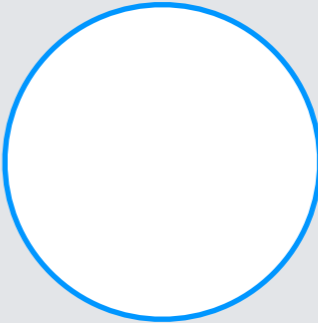


ALMA

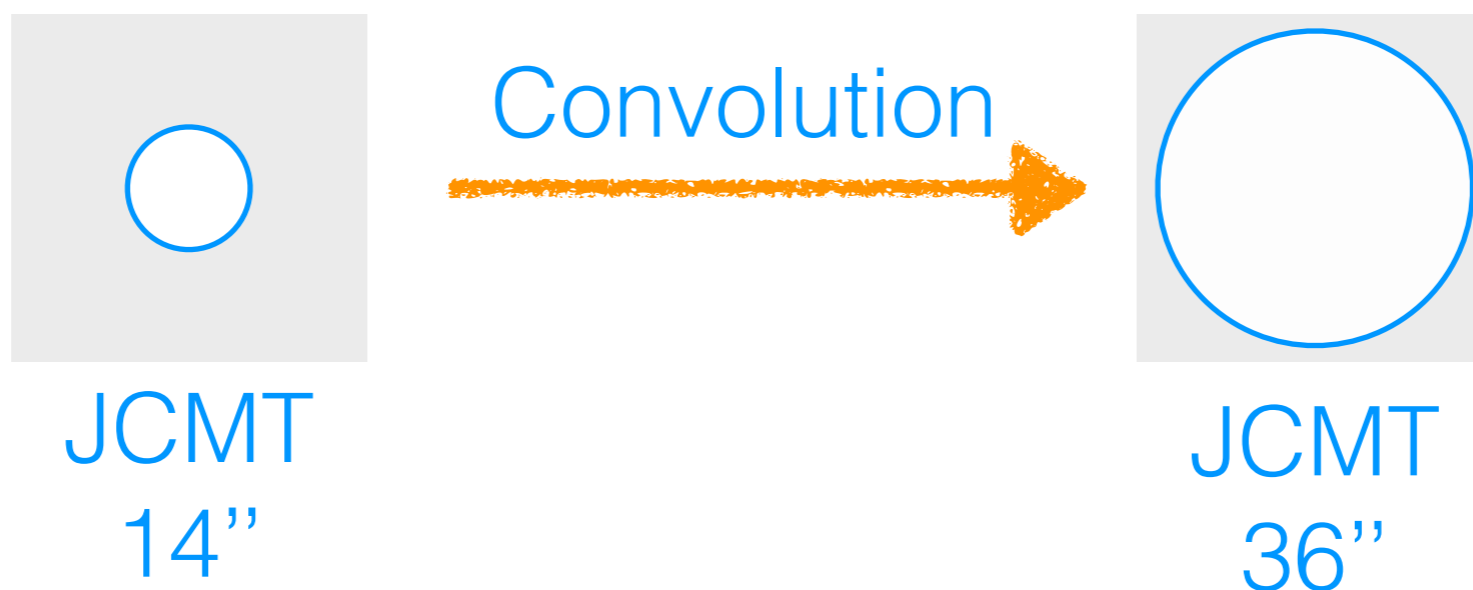


Herschel

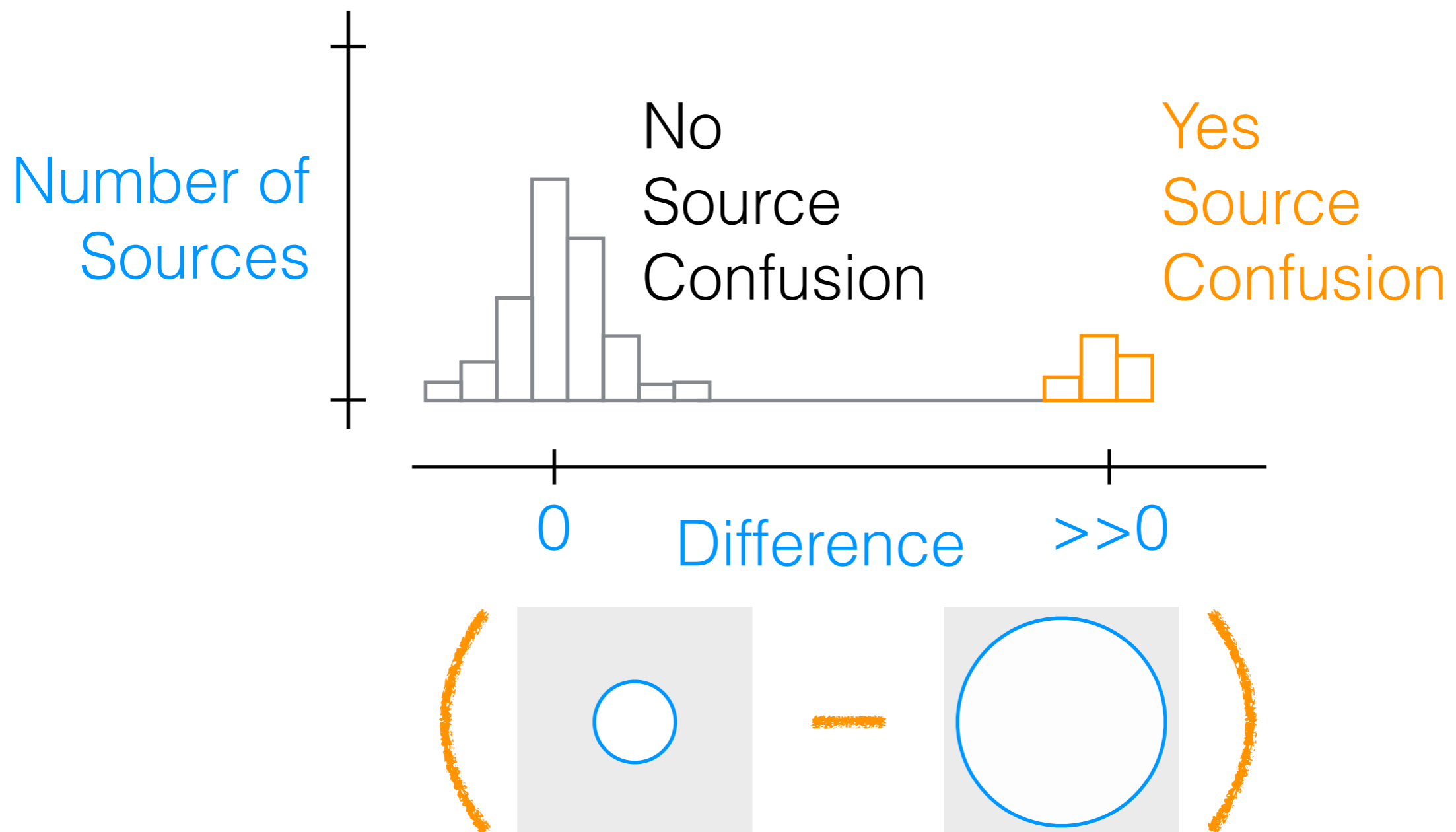
Source confusion effects need to be measured

	<i>Herschel</i>		
λ [μm]	250	350	500
Angular size	18"	25"	36"
Surface	158%	306%	634%
Beam size			

JCMT's luminosities at different resolutions estimate source confusion



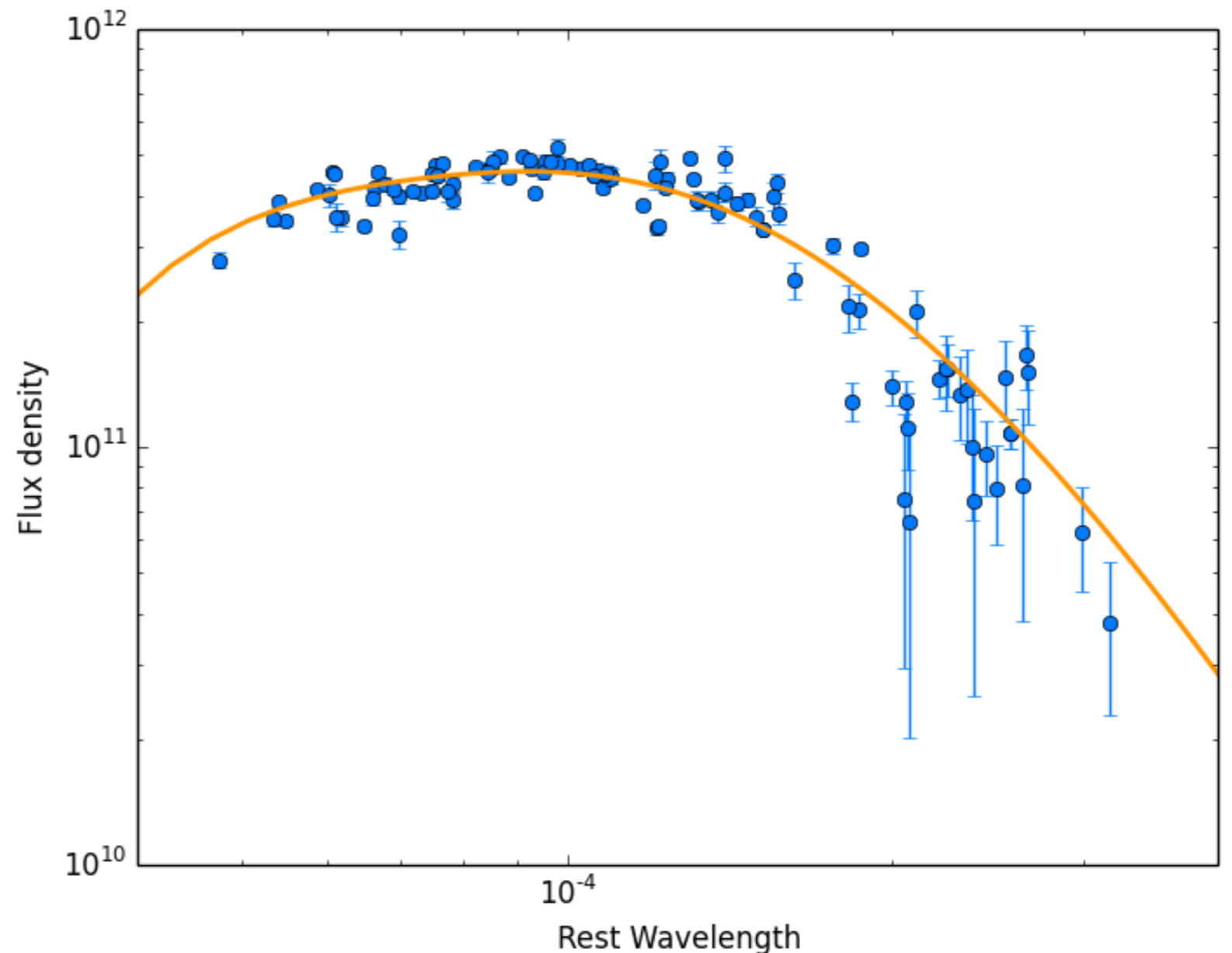
JCMT's luminosities at different resolutions estimate source confusion



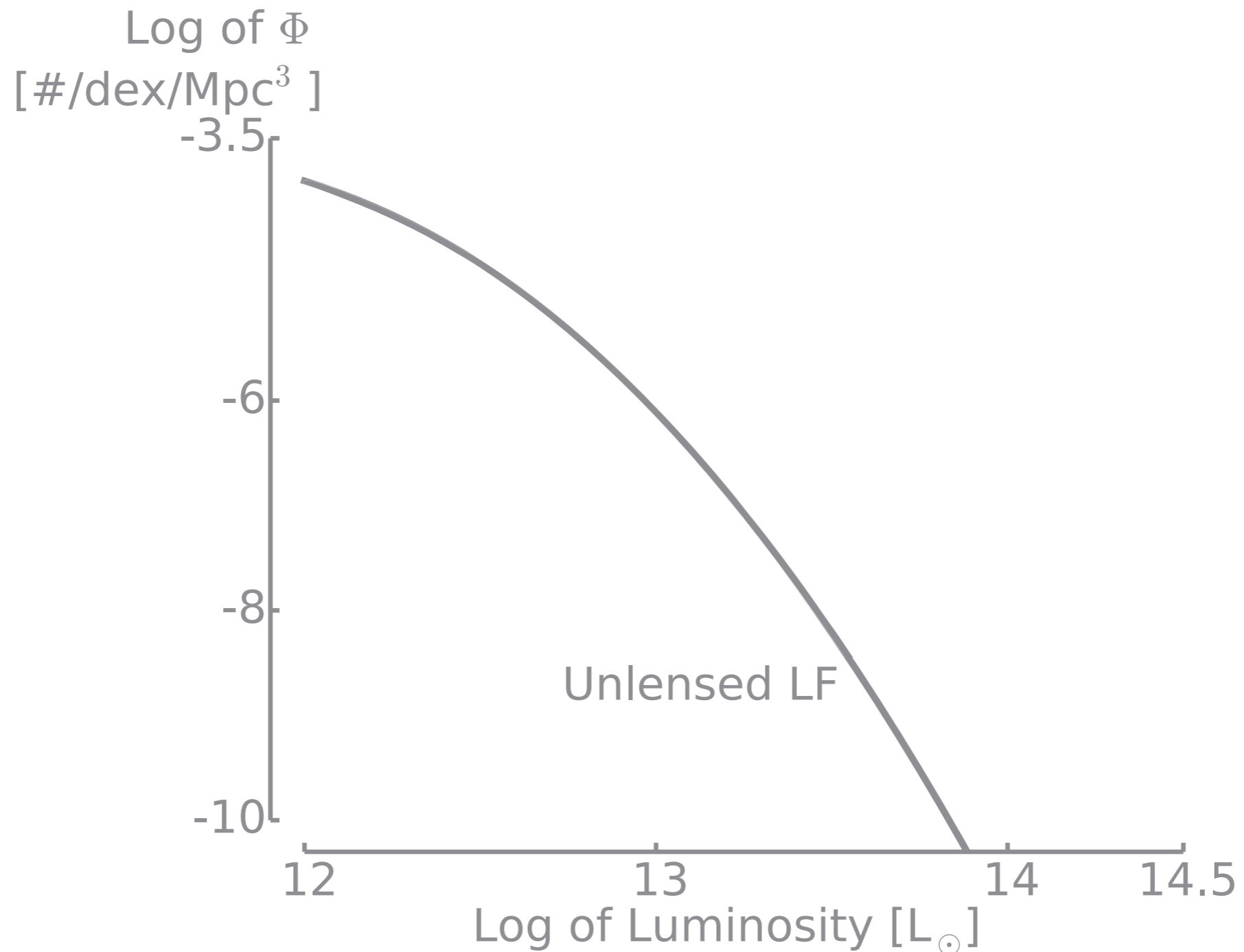
Template SED made from 26 spectroscopic sources

Two-temperature grey-body SED fit

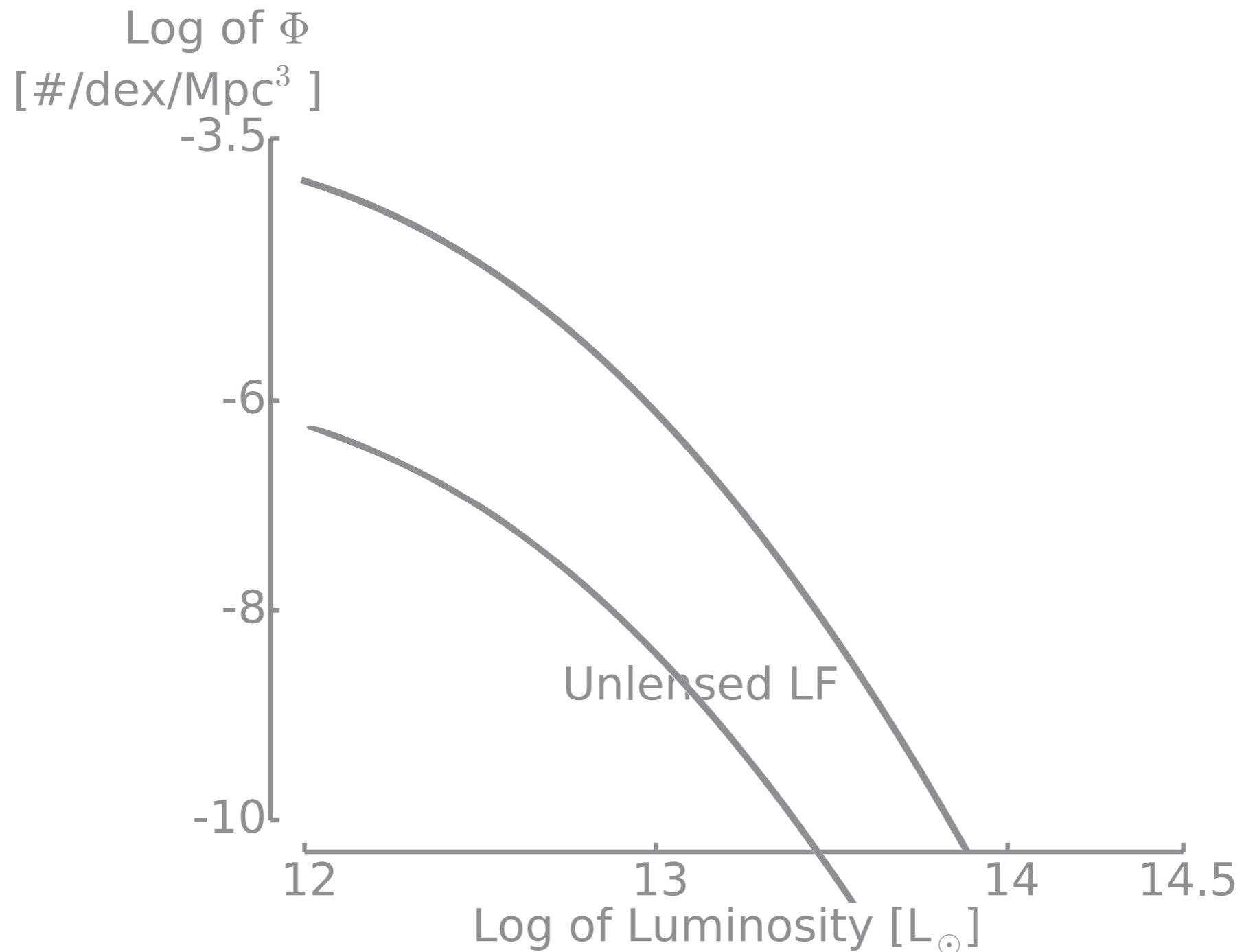
$$S_\nu = A \left[B_\nu(T_h) \nu^\beta + \alpha B_\nu(T_c) \nu^\beta \right]$$



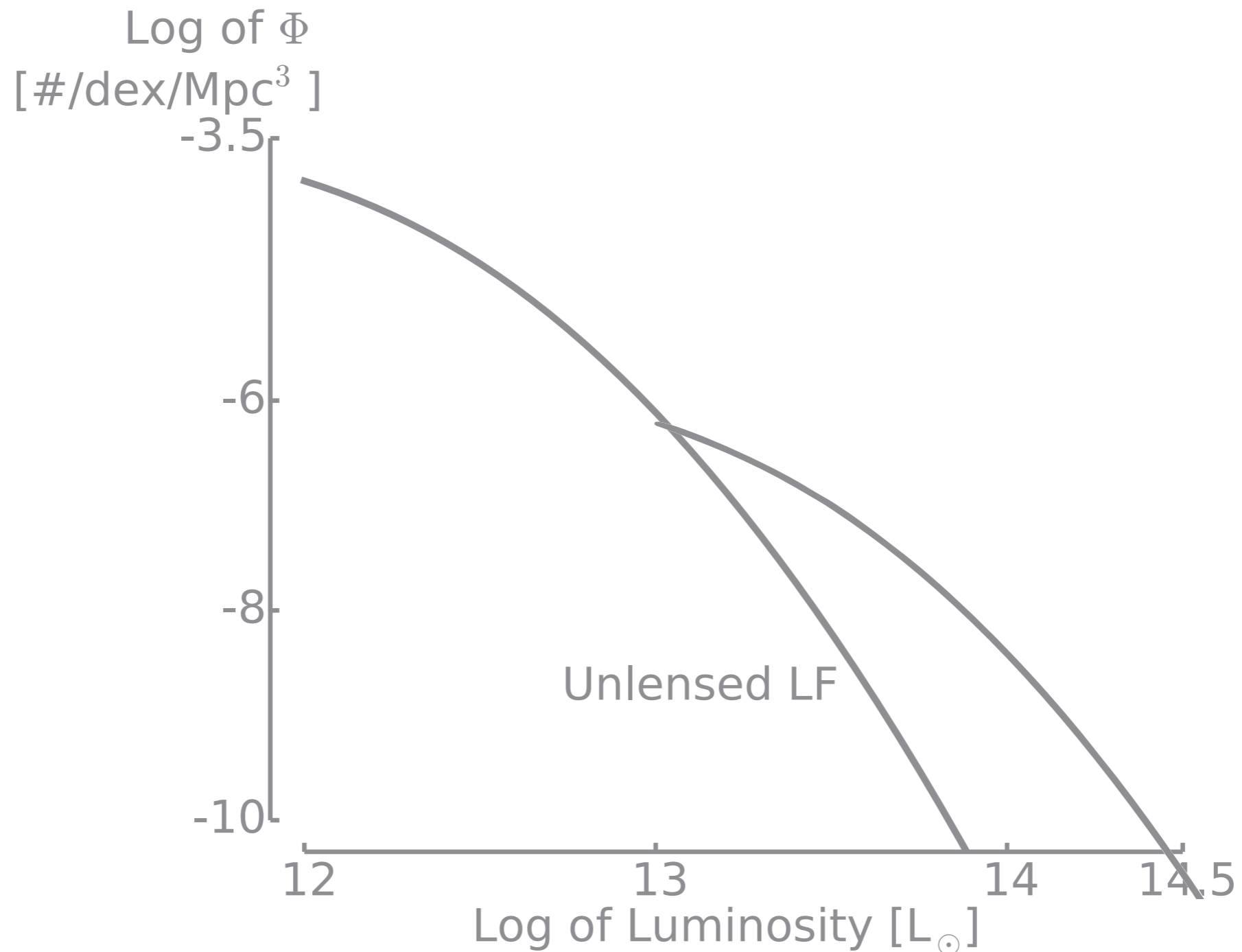
Luminosity functions provide global information on lensing



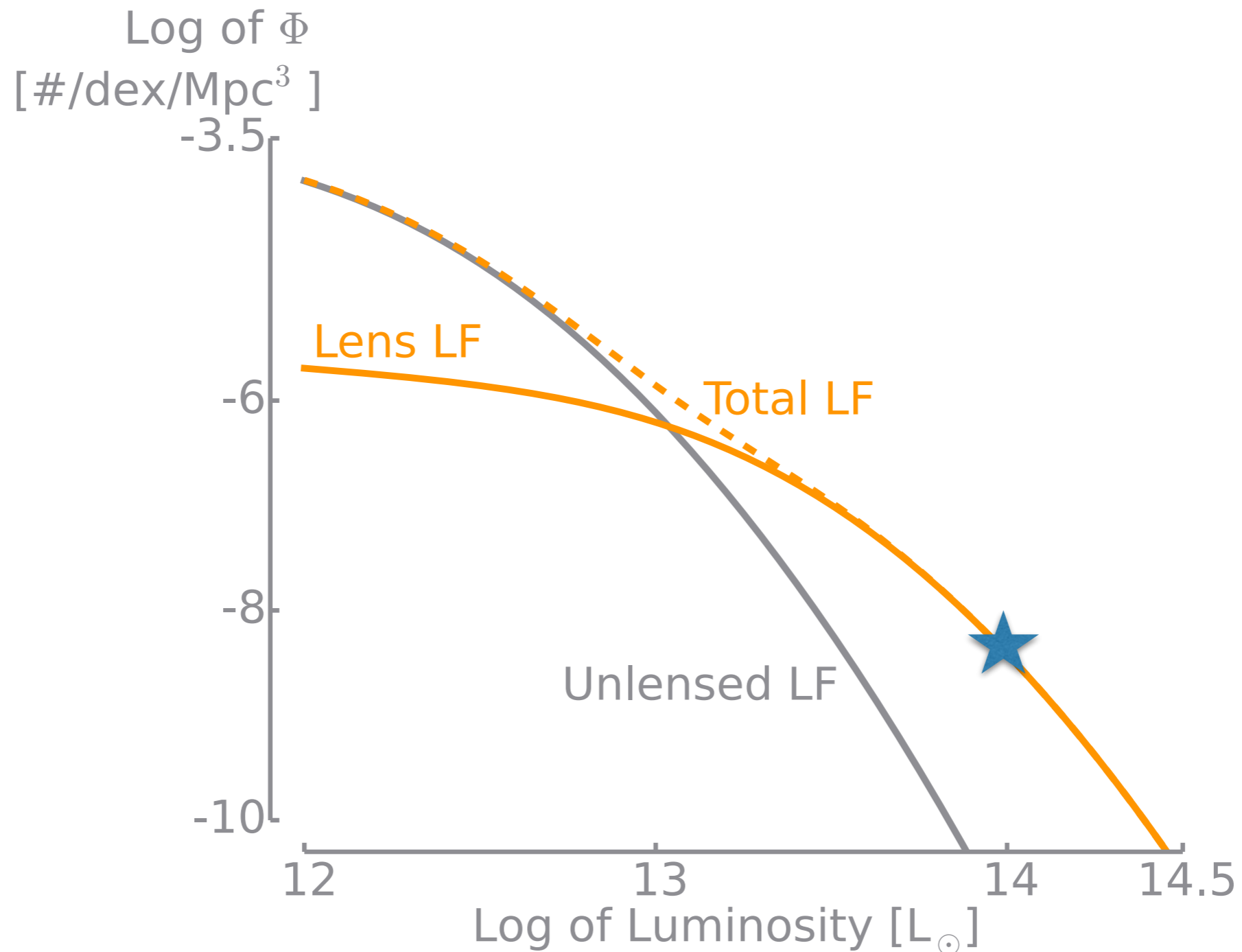
Luminosity functions provide global information on lensing



Luminosity functions provide global information on lensing

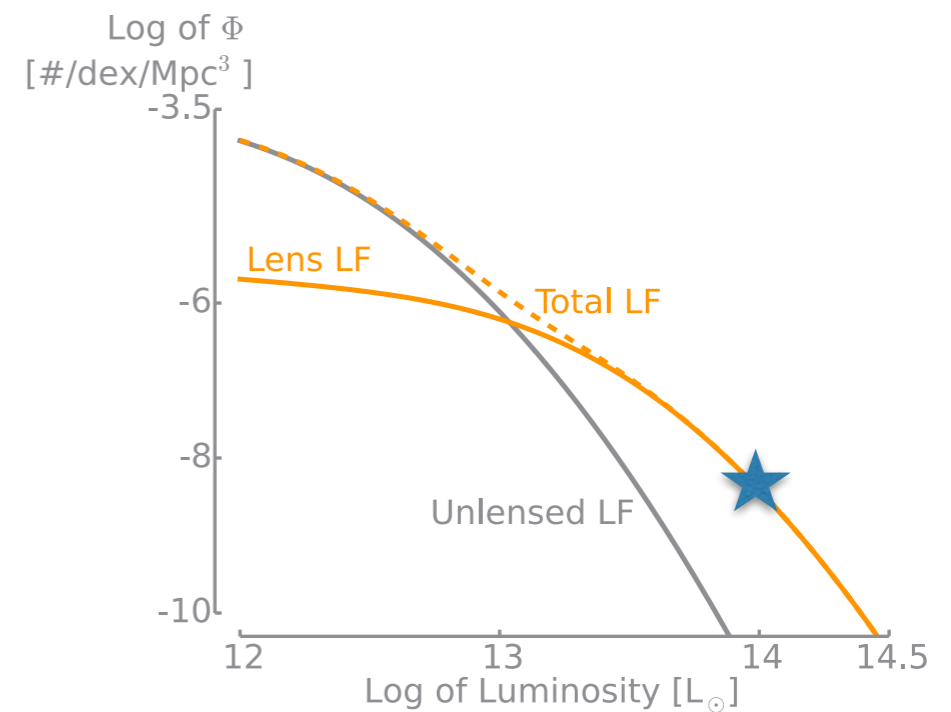
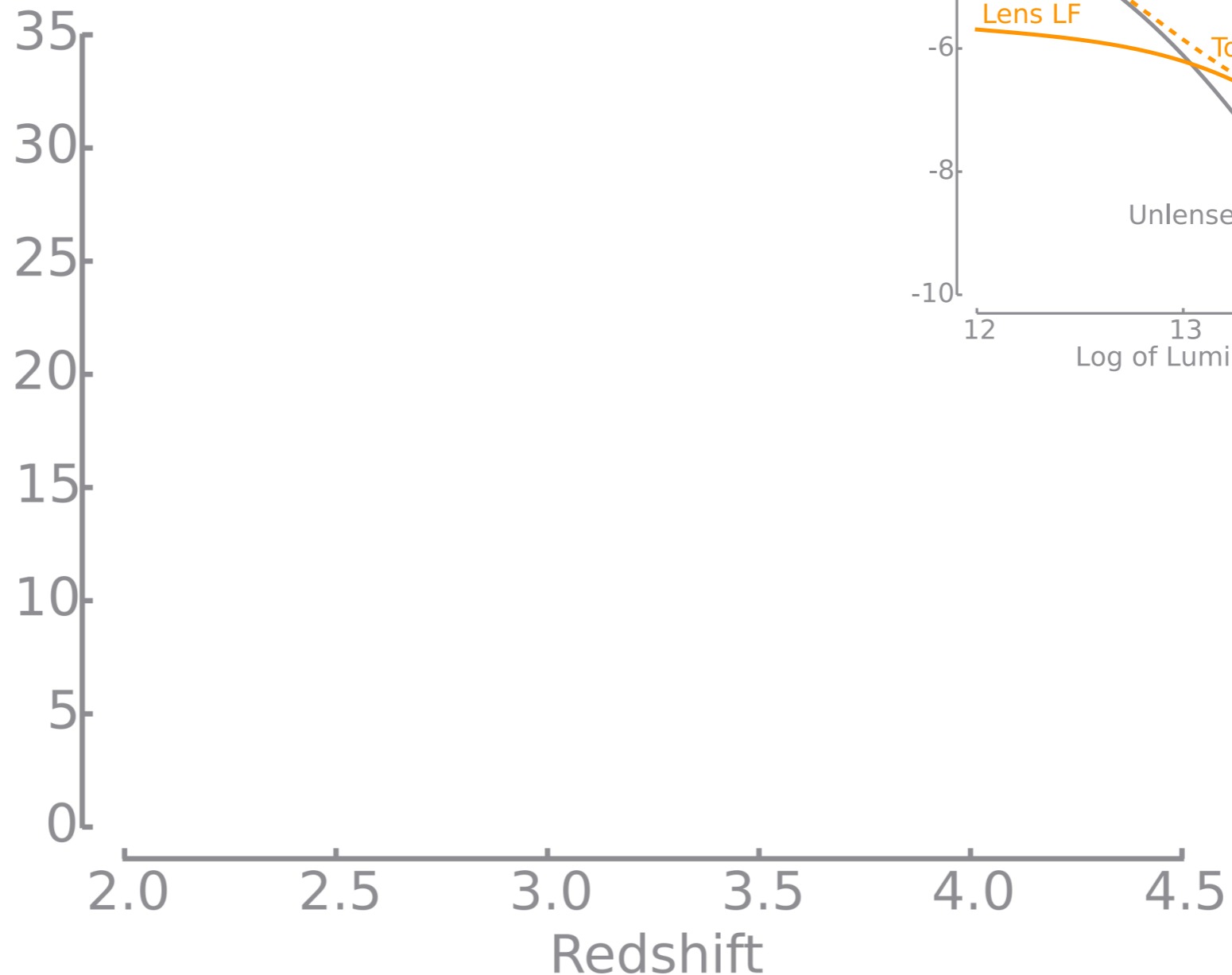


Luminosity functions provide global information on lensing

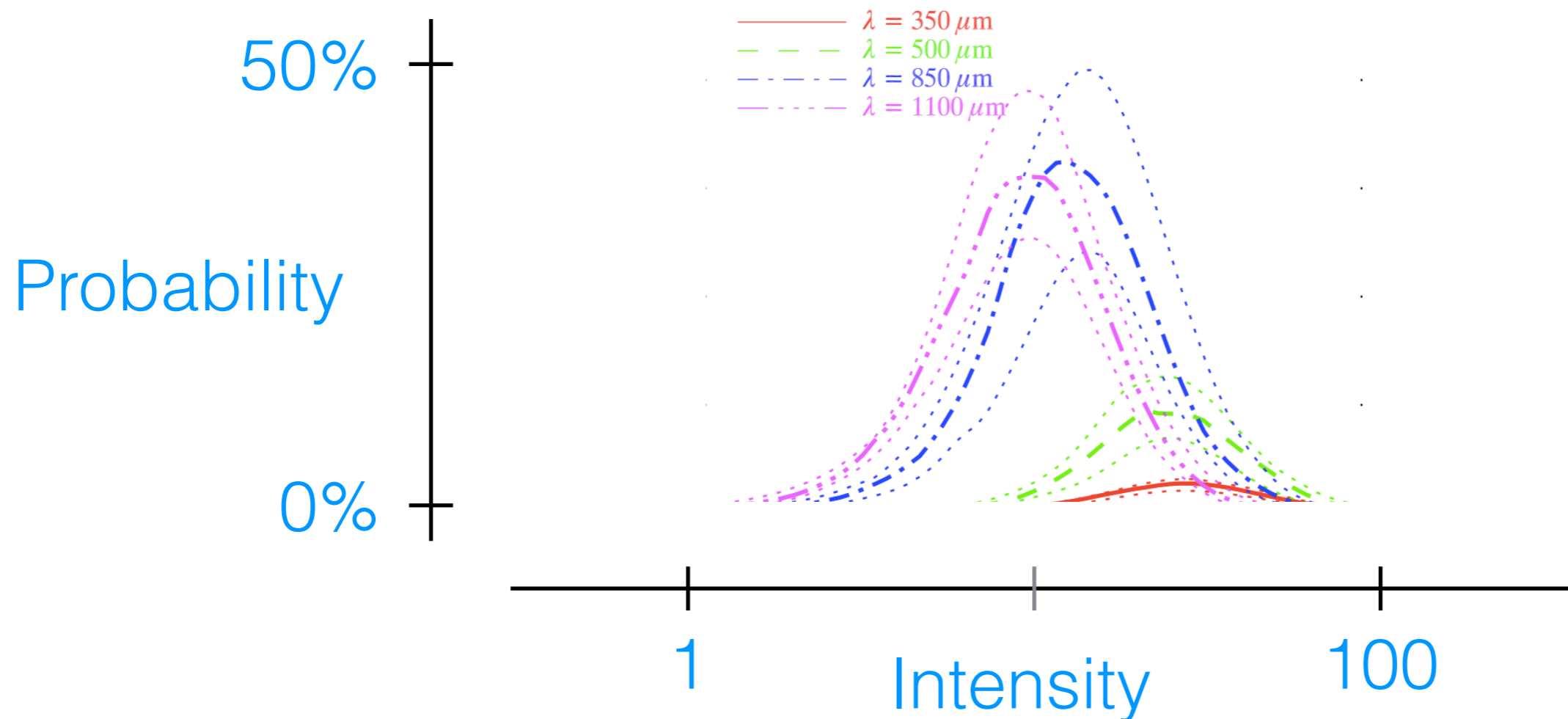
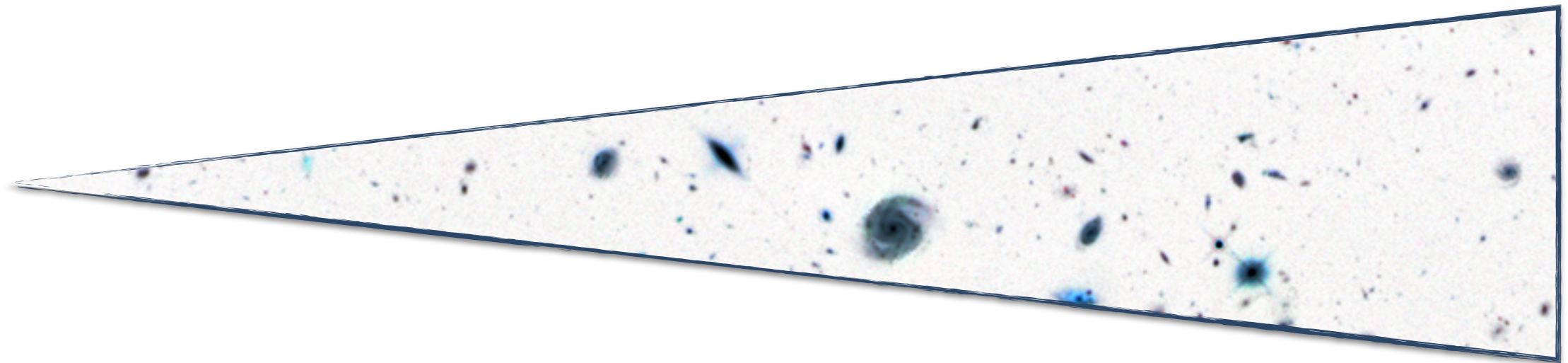


First analysis of the sources suggest a lensed sample bias

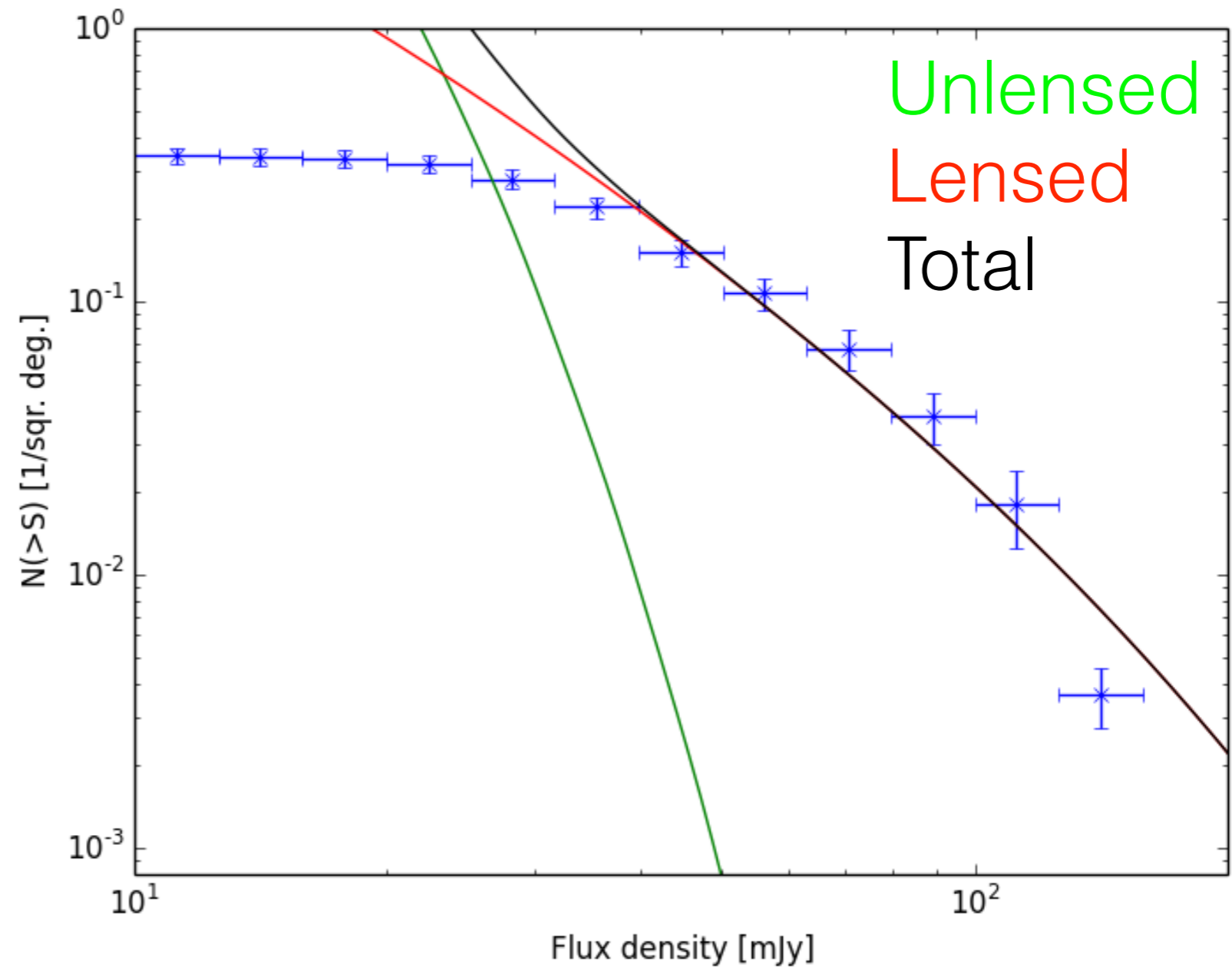
Galaxies



Cosmological models provide lensing probabilities



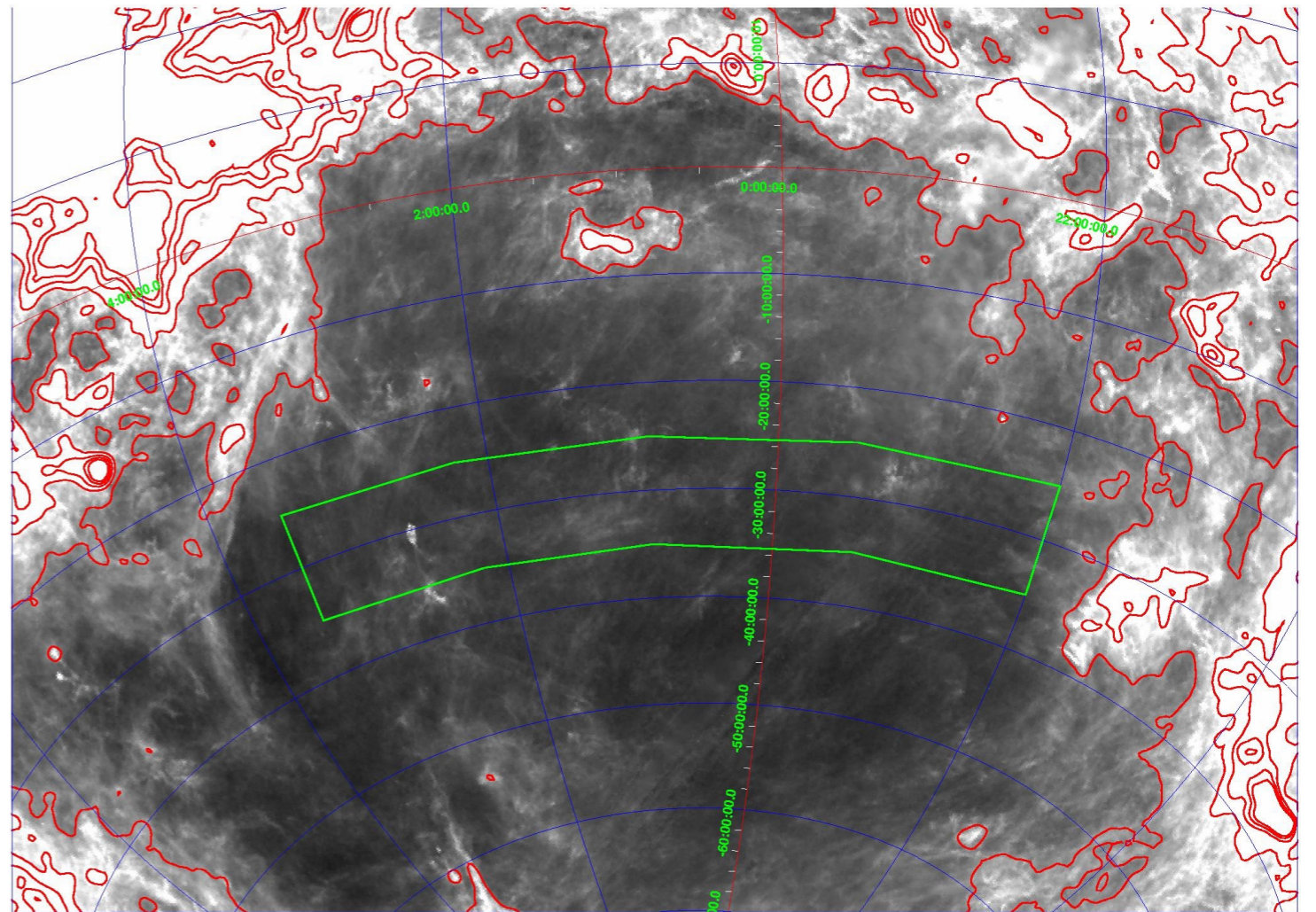
Preliminary results show a lensing-biased survey



Resolved images are necessary for definitive lensing evidence

KIDS and VIKINGS surveys will provide information on the lenses

This will help us improve lens-finding algorithms



Lens hunting with *Herschel*

Tom Bakx

